

900 feet high. Something similar is well known to the inhabitants of Arequipa, Peru. The city is built at the base of the extinct volcano "Misti," which rises above the plaza of Arequipa to a height of about 12,500 feet; Arequipa itself being over 7,000 feet above the level of the sea. It is not an uncommon occurrence (during the fall of the year, February and March) in the morning, from sunrise till about ten o'clock, to see a succession of clouds rolling along the summit from N.E. to S.W., much as if huge masses of white smoke were issuing from the extinct crater. These clouds are either suddenly shot upward by meeting the current from the S.W. and lost at a distance of from 30,000 to 40,000 feet to the eastward from the summit, or else, rolling over the summit, they are carried by the easterly breezes till they become absorbed by the dryer and warmer air of the region to the southward of Misti.

It must be remembered that between Arequipa and the sea, at a distance of not more than thirty miles, extends the great sandy desert of Islay, having an average breadth of about twenty-five miles, and before the days of the railroad the great terror of all travellers from the sea-coast to the interior. Of course the winds blowing across this desert (a part of the great rainless belt



of Peru) are greatly heated at all seasons of the year. The eastern slope of Misti, on the contrary, forms the edge of the elevated plateau extending for more than 150 miles to the eastern slope of the Andes, having an altitude of from 10,000 to 14,000 feet, and the amount of rain falling in this district is very great.

The formation of the cloud, seen from Arequipa on the summit only of Misti, is plainly seen from the railroad leading to Puno, which, after leaving Arequipa, makes a gigantic sweep northward round the Chacharni Mountains, and winds its way eastward behind Misti at a height of about 12,500 feet above the level of the sea. There I have several times seen masses of vapour, condensed into huge white clouds rolling along the slopes of Misti, travel up with great rapidity towards the summit, and either follow its crest as described above, or become at once reabsorbed on reaching the top. This shows plainly that the clouds seen from Arequipa are not due to volcanic action; the Indians also all agree in stating that there is no tradition among them of Misti having been active. I enclose a sketch of Misti and its cloud from a photograph obtained during my visit to Peru.

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(The Effect of Waves)

It is generally believed that at a moderate depth the influence of heavy waves ceases, and that during a hurricane all is quiet a few fathoms beneath the surface. If this be correct, why should a swell show such a marked increase in height when it rolls over the edge of soundings?

On the parallel of Cape Clear, in longitude 15° W., seamen are familiar with this phenomenon, although the depth is nearly

five hundred fathoms; at times it is so marked that the dead reckoning may be checked by carefully noting the increase in the depth of the hollow of the waves. Shortly after the edge of soundings is passed the sea becomes more regular, and consequently less dangerous to deeply laden vessels.

Anyone who has watched during a moderate breeze the commotion of the water close to a quay wall can form a good idea of the ocean when it receives its first check against the Irish Plateau; the great waves twist around each other, run up and down in heaps, and then fall suddenly as if bereft, in a great measure, of their forward motion.

Again, it is a well-known fact that during a "norther" in the Gulf of Mexico the frailest vessels weather out the storm if they can cross the edge of the Campeachy Banks; a striking proof that at a depth of over fifty fathoms there is sufficient abrasion to destroy the force of the heaviest wave in a very effectual style. On one occasion the writer witnessed this remarkable fact by running from a turbulent sea into comparative smooth water in this locality.

On George's Shoals, off Nantucket, during a heavy gale, the New York pilots and masters of coasting vessels assert that sand is frequently left on deck after a sea has broken on board, although the depth of water may be twelve or fourteen fathoms. It must require an enormous amount of ebullition at the bottom to raise such dense matter to the surface through such a distance; for a cubic foot of ordinary sea-sand weighs about 100 pounds.

In this wild spot the tide, which frequently runs with a velocity of three miles per hour, would assist the lifting power of the wave if running counter to it. During a winter gale, when the strong springs are thus running, the confusion of the sea is indescribable, although the depth may be thirty fathoms. The shortness of the sea (*i.e.* the distance between the crests of the waves) on the banks of Newfoundland, where the soundings are from thirty to fifty fathoms, is noticed by all the navigators of the Western Atlantic, as it reduces the speed of an ocean steamer more than the heavier waves of deeper water with a similar force of wind will do. It is evident that this can only arise from the friction of the bottom, as the waves increase in height when deeper water is reached a short distance to the eastward.

In the Gulf Stream north of the Straits of Bemine, after a "norther" has blown a few hours, the surface of the sea is covered with lanes of weed, although only a few patches might have been seen before the commencement of the gale. As these lanes are often at a considerable distance from shoal water, which lies at right angles to the direction of the current and wind, it is evident they must have grown near the spot where they float, and been torn from their moorings by the mechanical force of the waves.

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THE TOTAL SOLAR ECLIPSE OF 1605, Oct. 12.—Clavius, observing the solar eclipse of April 9, 1567, at its maximum, remarked "a narrow ring of light round the moon which he supposed to be the margin of the solar disc." Kepler, however, maintained that this could not be in reality a portion of the sun, because the moon's apparent diameter at the time must have been greater than that of the sun, and he concluded, as Prof. Grant relates in his "History of Physical Astronomy," that the sun must have been totally covered by the moon while the narrow ring of light was visible, a phenomenon again exhibited in the total eclipse of Oct. 12, 1605, which was observed at Naples. Of this eclipse Kepler says (*De Stella Nova in pede Serpentarii*)—"Accuratè rectum fuisse totum Solem, quod quidem non diu duraverit; in medio, ubi Luna, fuisse speciem quasi nigræ nubis; circumcirca rubentem et flammæ splendorem, æqualis undique latitudinis, qui bonam cœli partem occupaverit: E regioni Solis, versus Septentrionem, cœlum obscurum planè, et cum profunda nox est; stellas tamen non visas."

Adopting the same system of elements of the lunar motions, employed in previous calculations of past eclipses, the results of which have appeared in this column, we have the following elements of the eclipse to which Kepler refers:—

